

Package ‘UPMASK’

October 12, 2022

Type Package

Title Unsupervised Photometric Membership Assignment in Stellar Clusters

Version 1.2

Date 2019-01-28

Maintainer Alberto Krone-Martins <algol@sim.ul.pt>

Description An implementation of the UPMASK method for performing membership assignment in stellar clusters in R. It is prepared to use photometry and spatial positions, but it can take into account other types of data. The method is able to take into account arbitrary error models, and it is unsupervised, data-driven, physical-model-free and relies on as few assumptions as possible. The approach followed for membership assessment is based on an iterative process, dimensionality reduction, a clustering algorithm and a kernel density estimation.

Depends R (>= 3.0)

License GPL (>= 3)

Imports parallel, MASS, RSQLite, DBI, dimRed, loe

NeedsCompilation no

RoxygenNote 6.0.1

Author Alberto Krone-Martins [aut, cre],
Andre Moitinho [aut],
Eduardo Bezerra [ctb],
Leonardo Lima [ctb],
Tristan Cantat-Gaudin [ctb]

Repository CRAN

Date/Publication 2019-02-01 17:43:33 UTC

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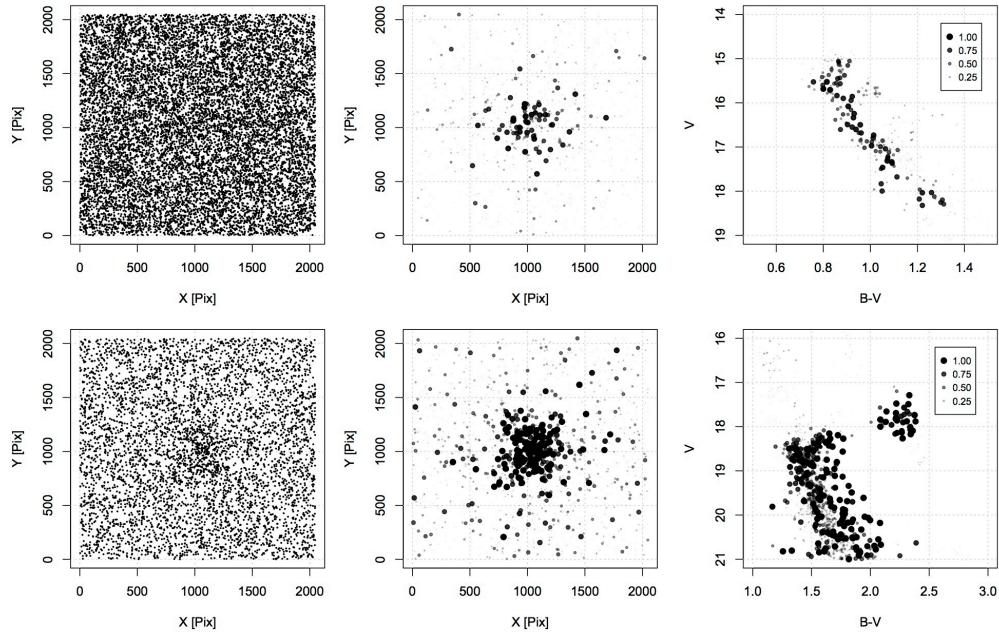
UPMASK-package

*Unsupervised Photometric Membership Assignment in Stellar Clusters***Description**

An implementation of the UPMASK method for performing membership assignment in stellar clusters in R. It is prepared to use photometry and spatial positions, but it can take into account other types of data. The method is able to take into account arbitrary error models, and it is unsupervised, data-driven, physical-model-free and relies on as few assumptions as possible. The approach followed for membership assessment is based on an iterative process, principal component analysis, a clustering algorithm and a kernel density estimation.

Details

Package:	UPMASK
Type:	Package
Version:	1.2
Date:	2017-06-09
License:	GPL (>= 3)



The two main functions in the UPMASK package are [UPMASKfile](#) and [UPMASKdata](#). The later will run the UPMASK method on data inside a data frame, while the former will perform the analysis on a file (the later deals with loading a file inside a data frame, calling the [UPMASKdata](#) function and writing the results to an output file).

The package includes data from two simulated fields comprising simulated data from cluster and field stars – to be used for demonstration. The analysis of these files using UPMASK lead to results presented in the figures above (from [Krone-Martins&Moitinho, 2014](#)), which show the spatial positions of the objects in the original datasets (in the left), the same objects but color coded by membership probability after the UPMASK analysis (in the center) and the color-magnitude diagram of all the stars also color coded by membership probability (in the right).

Author(s)

Alberto Krone-Martins, Andre Moitinho

Maintainer: Alberto Krone-Martins <algol@sim.ul.pt>

References

[Krone-Martins, A. & Moitinho, A., Astronomy&Astrophysics, v.561, p.A57, 2014](#)

See Also

[UPMASKfile](#), [UPMASKdata](#)

Examples

```
## Not run:
```

```

#
# Example of how to run UPMASK using data from a file
# Note: serious analysis require larger nRuns, and see UPMASKfile documentation
# for the parametrization.
# Write a string with the filename of the input and output files
inputFileName <- system.file("extdata",
                            "oc_12_500_1000_1.0_p019_0880_1_25km_120nR_withcolors.dat", package="UPMASK")
outputFileName <- file.path(tempdir(), "RESULTS.dat")
# Run UPMASK
UPMASKfile(inputFileName, outputFileName, nRuns=5, starsPerClust_kmeans=25,
            verbose=TRUE, fileWithHeader=TRUE)
# Done, the results are written to the file outputFileName

#
# Example of how to run UPMASK using data from a data frame
# Note: serious analysis require larger nRuns, and see UPMASKdata documentation
# for the parametrization.
# Load the data into a data frame
inputFileName <- system.file("extdata",
                            "oc_12_5000_4000_4.0_p019_0900_1_15km_120nR_withcolors.dat", package="UPMASK")
ocData <- read.table(inputFileName, header=TRUE)
# Run UPMASK
upmaskRes <- UPMASKdata(ocData, nRuns=5, starsPerClust_kmeans=25, verbose=TRUE)
# Done, the results are in the data frame upmaskRes

# Clean the environment
rm(list=c("inputFileName", "outputFileName", "ocData", "upmaskRes"))

## End(Not run)

```

analyse_randomKde2d *Perform analysis of random 2d distributions*

Description

analyse_randomKde2d will compute statistics from uniformly randomly created 2D fields based on Kernel Density Estimations (calling the code [create_randomKde2d](#)).

Usage

```
analyse_randomKde2d(nfields=100, nstars, maxX, maxY, nKde=50,
showStats=FALSE, returnStats=TRUE)
```

Arguments

nfields	an integer with the number of individual field realisations
nstars	an integer with the number of stars to consider
maxX	the length of the field in X
maxY	the length of the field in Y

nKde	the number of samplings of the kernel in each direction
showStats	a boolean indicating if the user wants to see statistics
returnStats	a boolean indicating if the user wants statistics to be returned

Value

A data frame with the mean and sd fields containing the results of the random field analysis.

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Runs the analysis on random fields
toyRes <- analyse_randomKde2d(100, 200, 100, 100, showStats=TRUE)

# Clean the environment
rm(toyRes)
```

analyse_randomKde2d_AutoCalibrated

Perform analysis of random 2d distributions (auto calibrated)

Description

The experimental code `analyse_randomKde2d_AutoCalibrated` computes statistics from uniformly randomly created 2D fields based on kernel density estimations (using [create_randomKde2d](#)). It runs for as many interations as necessary for the statistical measuremnt stabilize (but it will abort after `maxIter` iterations is reached).

Usage

```
analyse_randomKde2d_AutoCalibrated(nstars, maxX, maxY, nKde=50, maxIter=40,
                                     showStats=FALSE, returnStats=TRUE)
```

Arguments

nstars	an integer with the number of stars to consider
maxX	the length of the field in X
maxY	the length of the field in Y
nKde	the number of samplings of the kernel in each direction
maxIter	an integer with the maximum number of iterations
showStats	a boolean indicating if the user wants to see statistics
returnStats	a boolean indicating if the user wants statistics to be returned

Value

A data frame with the `mean` and `sd` fields containing the results of the random field analysis.

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Runs the analysis on random fields
toyRes <- analyse_randomKde2d_AutoCalibrated(100, 100, 100, 100, showStats=TRUE)

# Clean the environment
rm(toyRes)
```

`analyse_randomKde2d_smart`

Perform analysis of random 2d distributions

Description

`analyse_randomKde2d_smart` will compute statistics from uniformly randomly created 2D fields based on Kernel Density Estimations (calling the code `analyse_randomKde2d`). However, if a random field using the same number of stars was already computed in this run of UPMASK, it will avoid computing it again and will return the value that is stored in a SQLite database table. If the random field was not yet analysed, it will run the analysis, store the result in the database table, and return the value.

Usage

```
analyse_randomKde2d_smart(nfields=100, nstars, maxX, maxY, nKde=50,
showStats=FALSE, returnStats=TRUE, smartTableDB)
```

Arguments

<code>nfields</code>	an integer with the number of individual field realisations
<code>nstars</code>	an integer with the number of stars to consider
<code>maxX</code>	the length of the field in X
<code>maxY</code>	the length of the field in Y
<code>nKde</code>	the number of samplings of the kernel in each direction
<code>showStats</code>	a boolean indicating if the user wants to see statistics
<code>returnStats</code>	a boolean indicating if the user wants statistics to be returned
<code>smartTableDB</code>	a database connection to the smart look-up table

Value

A data frame with the `mean` and `sd` fields containing the results of the random field analysis.

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Create the smart look-up table
library(RSQLite)
stcon <- create_smartTable()

# Runs the analysis on random fields
system.time(
toyRes1 <- analyse_randomKde2d_smart(300, 200, 100, 100, smartTableDB=stcon)) # slow
system.time(
toyRes2 <- analyse_randomKde2d_smart(300, 200, 100, 100, smartTableDB=stcon)) # quick

# Clean the environment
rm(list=c("toyRes1", "toyRes2"))
dbDisconnect(stcon)
```

`create_randomKde2d`

Compute the density based distance quantity using a 2D Kernel Density Estimation

Description

`create_randomKde2d` will compute the 2D Kernel Density Estimation for a random sampling of the requested number of points and will return the quantity $(\max(d) - \text{mean}(d))/\text{sd}(d)$, if the option `returnDistance` is set to TRUE.

Usage

```
create_randomKde2d(nstars, maxX, maxY, nKde=50, printPlots=FALSE,
showStats=FALSE, returnDistance=FALSE)
```

Arguments

<code>nstars</code>	an integer with the number of stars to consider
<code>maxX</code>	the length of the field in X
<code>maxY</code>	the length of the field in Y
<code>nKde</code>	the number of samplings of the kernel in each direction
<code>printPlots</code>	a boolean indicating if the user wants to see plots
<code>showStats</code>	a boolean indicating if the user wants to see statistics
<code>returnDistance</code>	a boolean indicating if the user wants statistics to be returned

Value

A double representing the density based distance quantity.

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., *A&A*, v.561, p.A57, 2014

Examples

```
# Compute the density based distance quantity with random fields
distVar <- create_randomKde2d(100, 10, 10, showStats=FALSE,
                                printPlots=FALSE, returnDistance=TRUE)

# Clean the environment
rm(distVar)
```

create_smartTable *Create a look up table*

Description

`create_smartTable` will create a look up table for random field analysis inside an SQLite database. The table is automatically filled each time UPMASK calls the function [analyse_randomKde2d_smart](#).

Usage

```
create_smartTable()
```

Value

A data base connection to the SQLite database containing the smartTable.

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Create the table
library(RSQLite)
stcon <- create_smartTable()

# Clean the environment
dbDisconnect(stcon)
```

getStarsAtHighestDensityRegion

Perform cut in the membership list based on the 2D space distribution

Description

getStarsAtHighestDensityRegion will compute the 2D Kernel Density Estimation for the requested subset of data and will return only the stars in the most dense region.

Usage

```
getStarsAtHighestDensityRegion(ocdata_out, threshold=2, posIdx=c(1,2),  
plotAnalysis=FALSE, verbose=FALSE)
```

Arguments

ocdata_out	a data frame to use
threshold	a double with the thresholding level
posIdx	an array of integers indicating the columns of the data frame containing the spatial positions
plotAnalysis	a boolean indicating if the analysis should be plotted
verbose	a boolean indicating if the code should be verbose

Value

A data frame with the objects which were selected from ocdata_out

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Create a simple data set  
toyDataDF <- data.frame(x=runif(50, 0, 10), y=runif(50, 0, 10), resMclust.class=rep(1, 50))  
toyDataDF <- rbind(toyDataDF, data.frame(x=rnorm(50, 2, 3),  
                                         y=rnorm(50, 4, 3), resMclust.class=rep(1, 50)))  
  
# Perform the XY density based cut  
toyRes <- getStarsAtHighestDensityRegion(toyDataDF)  
  
# Clean the environment  
rm(list=c("toyDataDF", "toyRes"))
```

innerLoop*UPMASK inner loop*

Description

`innerLoop` executes the UPMASK method's inner loop and returns the stars which were considered as cluster member stars.

The `innerLoop` perform the PCA, runs the clustering algorithm and check for overdensities in the spatial distribution for the clustered stars in the PC space using a 2d kernel density estimation.

Usage

```
innerLoop(ocdata_full, ocdata, classAlgol="kmeans", autoThresholdLevel=3,
          autoThreshold=TRUE, iiter=0, plotIter=FALSE, verbosity=1, starsPerClust_kmeans=50,
          nstarts_kmeans=50, runId=0, autoCalibrated=FALSE, stopIfEmpty=FALSE,
          positionDataIndexes=c(1,2), smartTableDB, nDimsToKeep=4, dimRed="PCA", scale=TRUE)
```

Arguments

<code>ocdata_full</code>	a data frame with the data to perform the analysis
<code>ocdata</code>	a data frame with the data to consider in the PCA step
<code>classAlgol</code>	a string indicating the type of clustering algorithm to consider. Only k-means is implemented at this moment (defaults to kmeans)
<code>autoThresholdLevel</code>	an integer indicating the level for thresholding of the spatial distribution
<code>autoThreshold</code>	a boolean indicating if autoThresolding should be adopted (defaults to TRUE)
<code>iiter</code>	and integer indicating the number of the iteration (passed by the outerLoop)
<code>plotIter</code>	a boolean indicating if the user wants to see iteration plots (defaults to FALSE)
<code>verbosity</code>	a flag indicating the verbosity level: it can be 0 (no screen output at all), 1 (minimum), >=2 (all)
<code>starsPerClust_kmeans</code>	an integer with the average number of stars per k-means cluster
<code>nstarts_kmeans</code>	an integer the amount of random re-initializations of the k-means clustering method (usually it is not necessary to modify this)
<code>runId</code>	an integer greater than zero indicating the run Id (passed by the outerLoop)
<code>autoCalibrated</code>	a boolean indicating if the number of random field realizations for the clustering check in the position space should be autocalibrated (experimental code, defaults to FALSE).
<code>stopIfEmpty</code>	a boolean indicating if the code should completely stop if no spatial clustering is detected (defaults to FALSE)
<code>positionDataIndexes</code>	an array of integers indicating the columns of the data frame containing the spatial position measurements

smartTableDB	a database connection to the smart look-up table
nDimsToKeep	an integer with the number of dimensions to consider (defaults to 4)
dimRed	a string with the dimensionality reduction method to use (defaults to PCA. The only other options are LaplacianEigenmaps or None)
scale	a boolean indicating if the data should be scaled and centered

Value

A data frame with objects considered as members at this iteration.

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., A&A, v.561, p.A57, 2014

Examples

```
## Not run:
# Perform a one run of the innerLoop using a simulated open cluster with
# spatial and photometric data
# Load the data into a data frame
fileName <- "oc_12_500_1000_1.0_p019_0880_1_25km_120nR_withcolors.dat"
inputFileName <- system.file("extdata", fileName, package="UPMASK")
ocData <- read.table(inputFileName, header=TRUE)
ocData <- data.frame(ocData, id=(1:length(ocData[,1]))) # create an id

# Prepare the data to run the inner loop
posIdx <- c(1,2)
photIdx <- c(3,5,7,9,11,19,21,23,25,27)

# Create the look up table
library(RSQLite)
stcon <- create_smartTable()

# Run the inner loop
innerLoopRes <- innerLoop(ocData, ocData[,photIdx], autoThresholdLevel=1, verbosity=2,
                           starsPerClust_kmeans=25, positionDataIndexes=posIdx,
                           smartTableDB=stcon)

# Clean the environment
rm(list=c("inputFileName", "ocData", "posIdx", "photIdx", "innerLoopRes",
         "fileName"))
dbDisconnect(stcon)

## End(Not run)
```

kde2dForSubset	<i>Compute the density based distance quantity using a 2D Kernel Density Estimation</i>
----------------	---

Description

kde2dForSubset will compute the 2D Kernel Density Estimation for the requested subset of data and will return the quantiy $(\max(d) - \text{mean}(d)) / \text{sd}(d)$ if the option `returnDistance` is set to TRUE.

Usage

```
kde2dForSubset(df, setw=1, n=50, showStats=TRUE, printPlots=TRUE,
  returnDistance=FALSE, positionDataIndexes=c(1,2))
```

Arguments

<code>df</code>	a data frame to use
<code>setw</code>	an integer with the class of the stars to perform the analysis
<code>n</code>	the number of points in the regular grid of the density estimation
<code>showStats</code>	a boolean indicating if the user wants to see output statistics
<code>printPlots</code>	a boolean indicating if the user wants to see plots
<code>returnDistance</code>	a boolean indicating if the distance between the max and the mean in units of standard deviations should be returned
<code>positionDataIndexes</code>	an array of integers indicating the columns of the file containing the spatial position measurements

Value

A double representing the density based distance quantity.

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., A&A, v.561, p.A57, 2014

Examples

```
# Create a simple data set with the values and errors
toyDataDF <- data.frame(x=runif(50, 0, 10), y=runif(50, 0, 10), resMclust.class=rep(1, 50))

# Run the KDE 2D analysis for the required subset
disV <- kde2dForSubset(toyDataDF, showStats=FALSE, printPlots=FALSE, returnDistance=TRUE)

# Clean the environment
rm(list=c("toyDataDF", "disV"))
```

meanThreeSigRej

Perform cuts in the data

Description

`meanThreeSigRej` will perform an interative rejection using the mean and three sigma of the data. The function will compute the mean and standard deviation of the input vector, reject all entries lying farther than three-sigma, and iterate until the mean varies (fractionaly) by less than the tolerance value. The function will return a data frame with the mean, standard deviation value and the number of iterations until the convergence was reached.

Usage

```
meanThreeSigRej(vec, maxI, tolerance)
```

Arguments

vec	a vector with the data to consider
maxI	an integer with the maximum amount of iterations allowed
tolerance	a double with the tolerance value (as (old-new)/old)

Value

A data frame with the fields `mean`, `sd` (the standard deviation) and `convergenceAtIter` (the iteration where the convergence was reached).

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Create a simple data set with the values and errors
toyData <- c(rnorm(30, mean=0, sd=5), 10000, 1000)

# Call the function to perform cuts
toyDataItStat <- meanThreeSigRej(toyData)

cat(paste("True mean      = 0\n"))
cat(paste("Before rejection mean =", round(mean(toyData), 2), "\n"))
cat(paste("After rejection mean  =", round(toyDataItStat$mean, 2), "\n"))

# Clean the environment
rm(list=c("toyData", "toyDataItStat"))
```

outerLoop

UPMASK outer loop

Description

`outerLoop` executes the UPMASK method's outer loop on a data frame, and returns another data frame as output, with the id of the object and it's classification as a stellar cluster member or not.

The `outerLoop` perform cuts in the data if necessary (by calling [performCuts](#)), take errors in the data table into account if the user request (by calling [innerLoop](#)), runs the inner loop (by calling [innerLoop](#)) until convergence of the membership list or util the maximum number of iterations is reached.

Usage

```
outerLoop(ocdata_full, positionDataIndexes=c(1,2),
photometricDataIndexes=c(3,5,7,9,11,19,21,23,25,27),
photometricErrorDataIndexes=c(4,6,8,10,12,20,22,24,26,28), threshold=1, maxIter=25,
plotIter=FALSE, verbose=FALSE, starsPerClust_kmeans=50, nstarts_kmeans=50,
finalXYCut=FALSE, autoCalibrated=FALSE, considerErrors=FALSE, run=0,
smartTableDB, nDimsToKeep=4, dimRed="PCA", scale=TRUE)
```

Arguments

<code>ocdata_full</code>	a data frame with the data to perform the analysis
<code>positionDataIndexes</code>	an array of integers indicating the columns of the data frame containing the spatial position measurements
<code>photometricDataIndexes</code>	an array of integers with the column numbers containing photometric measurements (or any other measurement to go into the PCA step)

photometricErrorDataIndexes	an array of integers with the column numbers containing the errors of the photometric measurements
threshold	a double indicating the thresholding level for the random field analysis
maxIter	an integer the maximum amount of iterations of the outer loop before giving up convergence (usually it is not necessary to modify this)
plotIter	a boolean indicating if the user wants to see iteration plots
verbose	a boolean indicating if the output to screen should be verbose
starsPerClust_kmeans	an integer with the average number of stars per k-means cluster
nstarts_kmeans	an integer the amount of random re-initializations of the k-means clustering method (usually it is not necessary to modify this)
finalXYCut	a boolean indicating if a final cut in the XY space should be performed (defaults to FALSE)
autoCalibrated	a boolean indicating if the number of random field realizations for the clustering check in the position space should be autocalibrated (experimental code, defaults to FALSE).
considerErrors	a boolean indicating if the errors should be taken into account
run	an integer greater than zero indicating the run number
smartTableDB	a database connection to the smart look-up table
nDimsToKeep	an integer with the number of dimensions to consider (defaults to 4)
dimRed	a string with the dimensionality reduction method to use (defaults to PCA. The only other options are LaplacianEigenmaps or None)
scale	a boolean indicating if the data should be scaled and centered

Value

A data frame with the id and class (member / not member) of each object at this run.

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., A&A, v.561, p.A57, 2014

Examples

```
## Not run:
# Perform a one run of the outerLoop using a simulated open cluster with
# spatial and photometric data
# Load the data into a data frame
fileName <- "oc_12_500_1000_1.0_p019_0880_1_25km_120nR_withcolors.dat"
inputFileName <- system.file("extdata", fileName, package="UPMASK")
ocData <- read.table(inputFileName, header=TRUE)
```

```

# Create the look up table
library(RSQLite)
stcon <- create_smartTable()

# Run the outer loop
posIdx <- c(1,2)
photIdx <- c(3,5,7,9,11,19,21,23,25,27)
photErrIdx <- c(4,6,8,10,12,20,22,24,26,28)
outerLoopRes <- outerLoop(ocData, posIdx, photIdx, PhotErrIdx,
                           starsPerClust_kmeans=25, verbose=TRUE, smartTableDB=stcon)

# Clean the environment
rm(list=c("inputFileName", "ocData", "posIdx", "photIdx", "photErrIdx",
         "outerLoopRes", "fileName"))
dbDisconnect(stcon)

## End(Not run)

```

performCuts*Perform cuts in the data***Description**

`performCuts` will perform cuts in the data. This function is provided as a place holder, and it is empty, but it is called by UPMASK, so if the user needs to perform cuts in the data for the UPMASK analysis, this function should be tailored.

Usage

```
performCuts(originalData)
```

Arguments

`originalData` a data frame to use as the baseline

Value

A data frame.

Author(s)

Alberto Krone-Martins, Andre Moitinho

Examples

```
# Create a simple data set with the values and errors
toyDataDF <- data.frame(x=runif(10, 0, 10), dx=rep(0.2, 10), y=runif(10, 0, 10),
                         dy=rep(0.1, 10))

# Call the function to perform cuts
newToyDataDF <- performCuts(toyDataDF)

# Clean the environment
rm(list=c("toyDataDF", "newToyDataDF"))
```

takeErrorsIntoAccount *Take Errors Into Account for UPMASK analysis*

Description

Based on a data frame containing measurements and errors, the `takeErrorsIntoAccount` will produce another data frame where each measurement of the original data frame is replaced by another value taken from a random distribution. The implemented error model is gaussian, so each value of the output data frame will be a random sampling from a gaussian distribution where the mean is the value in the original data frame (indicated by the `photometricDataIndexes` column argument) and the standard deviation is the value from its corresponding error (indicated by the `photometricErrorDataIndexes` column argument). The newly constructed dataframe is returned by the function.

The user can adapt this function so it can take any error model into account during the UPMASK analysis.

Usage

```
takeErrorsIntoAccount(originalData, dataIndexes, errorIndexes)
```

Arguments

- `originalData` a data frame to use as the baseline
- `dataIndexes` an array of integers indicating the columns corresponding to the measurements
- `errorIndexes` an array of integers indicating the columns corresponding to the errors

Value

A data frame with the new values sampled from the error distributions.

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., A&A, v.561, p.A57, 2014

Examples

```
# Create a simple data set with the values and errors
toyDataDF <- data.frame(x=runif(10, 0, 10), dx=rep(0.2, 10), y=runif(10, 0, 10),
                         dy=rep(0.1, 10))

# Apply the error models to create another data frame
newToyDataDF <- takeErrorsIntoAccount(toyDataDF, c(1,3), c(2,4))

# Plot the results
plot(toyDataDF$x, toyDataDF$y)
points(newToyDataDF$x, newToyDataDF$y, pch=19, cex=0.8, col="red")

# Clean the environment
rm(list=c("toyDataDF", "newToyDataDF"))
```

Description

UPMASKdata executes the UPMASK method on a data frame, and returns another data frame as output, including the membership analysis result as additional columns.

UPMASKdata is a method for performing membership assignment in stellar clusters. The distributed code is prepared to use photometry and spatial positions, but it can take into account other types of data as well. The method is able to take into account arbitrary error models (the user must rewrite the `takeErrorsIntoAccount` function), and it is unsupervised, data-driven, physical-model-free and relies on as few assumptions as possible. The approach followed for membership assessment is based on an iterative process, dimensionality reduction, a clustering algorithm and a kernel density estimation.

Usage

```
UPMASKdata(dataTable, positionDataIndexes=c(1,2),
photometricDataIndexes=c(3,5,7,9,11,19,21,23,25,27),
photometricErrorDataIndexes=c(4,6,8,10,12,20,22,24,26,28), threshold=1,
classAlgol="kmeans", maxIter=25, starsPerClust_kmeans=25, nstarts_kmeans=50,
nRuns=8, runInParallel=FALSE, paralelization="multicore", independent=TRUE,
verbose=FALSE, autoCalibrated=FALSE, considerErrors=FALSE,
finalXYCut=FALSE, nDimsToKeep=4, dimRed="PCA", scale=TRUE)
```

Arguments

dataTable	a data frame with the data to perform the analysis
positionDataIndexes	an array of integers indicating the columns of the data frame containing the spatial position measurements
photometricDataIndexes	an array of integers with the column numbers containing photometric measurements (or any other measurement to go into the PCA step)
photometricErrorDataIndexes	an array of integers with the column numbers containing the errors of the photometric measurements
threshold	a double indicating the thresholding level for the random field analysis
classAlgol	a string indicating the type of clustering algorithm to consider. Only k-means is implemented at this moment (defaults to kmeans)
maxIter	an integer the maximum amount of iterations of the outer loop before giving up convergence (usually it is not necessary to modify this)
starsPerClust_kmeans	an integer with the average number of stars per k-means cluster
nstarts_kmeans	an integer the amount of random re-initializations of the k-means clustering method (usually it is not necessary to modify this)
nRuns	the total number of individual runs to execute the total number of outer loop runs to execute
runInParallel	a boolean indicating if the code should run in parallel
parallelization	a string with the type of parallelization to use. the parallelization can be: "multicore" or "MPIcluster". At this moment only "multicore" is implemented (defaults to multicore).
independent	a boolean indicating if non-parallel runs should be completely independent
verbose	a boolean indicating if the output to screen should be verbose
autoCalibrated	a boolean indicating if the number of random field realizations for the clustering check in the position space should be autocalibrated (experimental code, defaults to FALSE).
considerErrors	a boolean indicating if the errors should be taken into account
finalXYCut	a boolean indicating if a final cut in the XY space should be performed (defaults to FALSE)
nDimsToKeep	an integer with the number of dimensions to consider (defaults to 4)
dimRed	a string with the dimensionality reduction method to use (defaults to PCA. The only other options are LaplacianEigenmaps or None)
scale	a boolean indicating if the data should be scaled and centered

Value

A data frame with the original data used to run the method and additional columns indicating the classification at each run, as well as a membership probability in the frequentist sense.

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., *A&A*, v.561, p.A57, 2014

Examples

```
## Not run:
# Analyse a simulated open cluster using spatial and photometric data
# Load the data into a data frame
fileNameI <- "oc_12_500_1000_1.0_p019_0880_1_25km_120nR_withcolors.dat"
inputFileName <- system.file("extdata", fileNameI, package="UPMASK")
ocData <- read.table(inputFileName, header=TRUE)

# Example of how to run UPMASK using data from a data frame
# (serious analysis require at least larger nRuns)
posIdx <- c(1,2)
photIdx <- c(3,5,7,9,11,19,21,23,25,27)
photErrIdx <- c(4,6,8,10,12,20,22,24,26,28)

upmaskRes <- UPMASKdata(ocData, posIdx, photIdx, PhotErrIdx, nRuns=2,
                         starsPerClust_kmeans=25, verbose=TRUE)

# Create a simple raw plot to see the results
pCols <- upmaskRes[,length(upmaskRes)]/max(upmaskRes[,length(upmaskRes)])
plot(upmaskRes[,1], upmaskRes[,2], col=rgb(0,0,0,pCols), cex=0.5, pch=19)

# Clean the environment
rm(list=c("inputFileName", "ocData", "posIdx", "photIdx", "photErrIdx",
         "upmaskRes", "pCols"))

## End(Not run)
```

Description

UPMASKfile executes the UPMASK method using a file as an input and writes another file as an output. This is a wrapper function that only reads a file into an R data frame, calls the **UPMASKdata** function using this data frame and the parameters passed by the user and writes the output into another file.

Usage

```
UPMASKfile(filenameWithPathInput, filenameWithPathOutput,
positionDataIndexes=c(1,2), photometricDataIndexes=c(3,5,7,9,11,19,21,23,25,27),
photometricErrorDataIndexes=c(4,6,8,10,12,20,22,24,26,28), threshold=1,
maxIter=20, starsPerClust_kmeans=50, nstarts_kmeans=50, nRuns=5,
runInParallel=FALSE, parallelization="multicore", independent=TRUE, verbose=FALSE,
autoCalibrated=FALSE, considerErrors=FALSE, finalXYCut=FALSE,
fileWithHeader=FALSE, nDimsToKeep=4, dimRed="PCA", scale=TRUE)
```

Arguments

filenameWithPathInput
 a string indicating the file containing the data to run UPMASK on (with full path)

filenameWithPathOutput
 a string indicating the file where the output shall be written (with full path)

positionDataIndexes
 an array of integers indicating the columns of the file containing the spatial position measurements

photometricDataIndexes
 an array of integers with the column numbers containing photometric measurements (or any other measurement to go into the PCA step)

photometricErrorDataIndexes
 an array of integers with the column numbers containing the errors of the photometric measurements

threshold a double indicating the thresholding level for the random field analysis

maxIter an integer the maximum amount of iterations of the outer loop before giving up convergence (usually it is not necessary to modify this)

starsPerClust_kmeans
 an integer with the average number of stars per k-means cluster

nstarts_kmeans an integer the amount of random re-initializations of the k-means clustering method (usually it is not necessary to modify this)

nRuns the total number of individual runs to execute the total number of outer loop runs to execute

runInParallel a boolean indicating if the code should run in parallel

parallelization a string with the type of parallelization to use. the parallelization can be: "multicore" or "MPIcluster". At this moment only "multicore" is implemented (defaults to multicore).

independent a boolean indicating if non-parallel runs should be completely independent

verbose a boolean indicating if the output to screen should be verbose

autoCalibrated a boolean indicating if the number of random field realizations for the clustering check in the position space should be autocalibrated (experimental code, defaults to FALSE).

considerErrors a boolean indicating if the errors should be taken into account

<code>finalXYCut</code>	a boolean indicating if a final cut in the XY space should be performed (defaults to FALSE)
<code>fileWithHeader</code>	a boolean indicating if the input file has a text header
<code>nDimsToKeep</code>	an integer with the number of dimensions to consider (defaults to 4)
<code>dimRed</code>	a string with the dimensionality reduction method to use (defaults to PCA. The only other options are LaplacianEigenmaps or None)
<code>scale</code>	a boolean indicating if the data should be scaled and centered

Author(s)

Alberto Krone-Martins, Andre Moitinho

References

Krone-Martins, A. & Moitinho, A., *A&A*, v.561, p.A57, 2014

Examples

```
## Not run:
# Analyse a simulated open cluster using spatial and photometric data
# Create strings with filenames
fileNameI <- "oc_12_500_1000_1.0_p019_0880_1_25km_120nR_withcolors.dat"
inputFileName <- system.file("extdata", fileNameI, package="UPMASK")
outputFileName <- file.path(tempdir(), "up-RESULTS.dat")

# Example of how to run UPMASK using data from a file
# (serious analysis require at least larger nRuns)
posIdx <- c(1,2)
photIdx <- c(3,5,7,9,11,19,21,23,25,27)
photErrIdx <- c(4,6,8,10,12,20,22,24,26,28)
UPMASKfile(inputFileName, outputFileName, posIdx, photIdx, photErrIdx, nRuns=5,
           starsPerClust_kmeans=25, verbose=TRUE, fileWithHeader=TRUE)

# Open the resulting file to inspect the results
tempResults <- read.table(outputFileName, header=TRUE)

# Create a simple raw plot to see the results
pCols <- tempResults[,length(tempResults)]/max(tempResults[,length(tempResults)])
plot(tempResults[,1], tempResults[,2], col=rgb(0,0,0,pCols), cex=0.5, pch=19)

# Clean the environment
rm(list=c("tempResults", "inputFileName", "outputFileName", "pCols", "fileNameI"))

## End(Not run)
```

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